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How does Microwave relate to other spectroscopies

- Different types of motion
 - Translational
 - Vibrational
 - Rotational.

What is Microwave Spectroscopy?

- Microwave stimulates Rotational translations
- Measures the rotational states of molecules
- *Gas Phase*
- Must have a dipole.

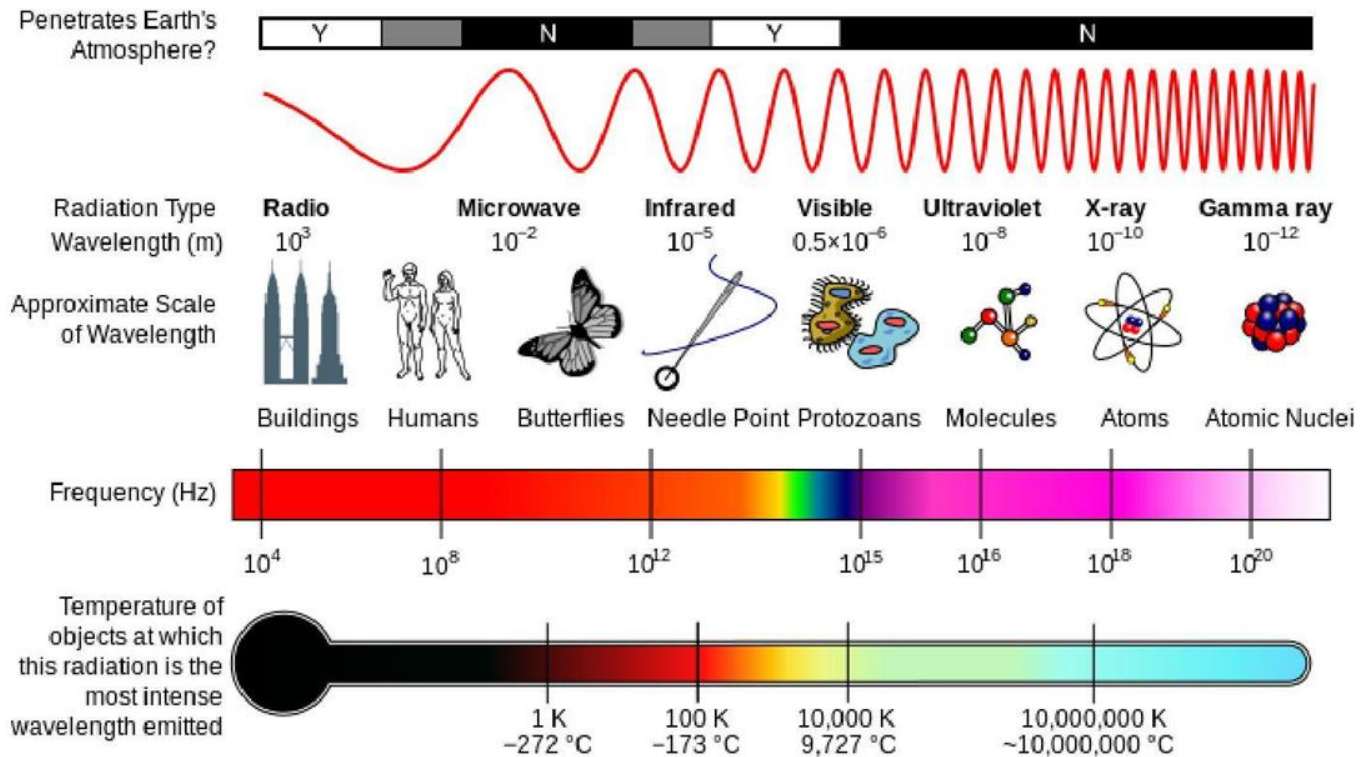
Applications of MW

- Measurement of bond lengths
- Observation by radio telescopes for life precursors in interstellar clouds
- Precise observation of translating stereochemistries and confirmation verification

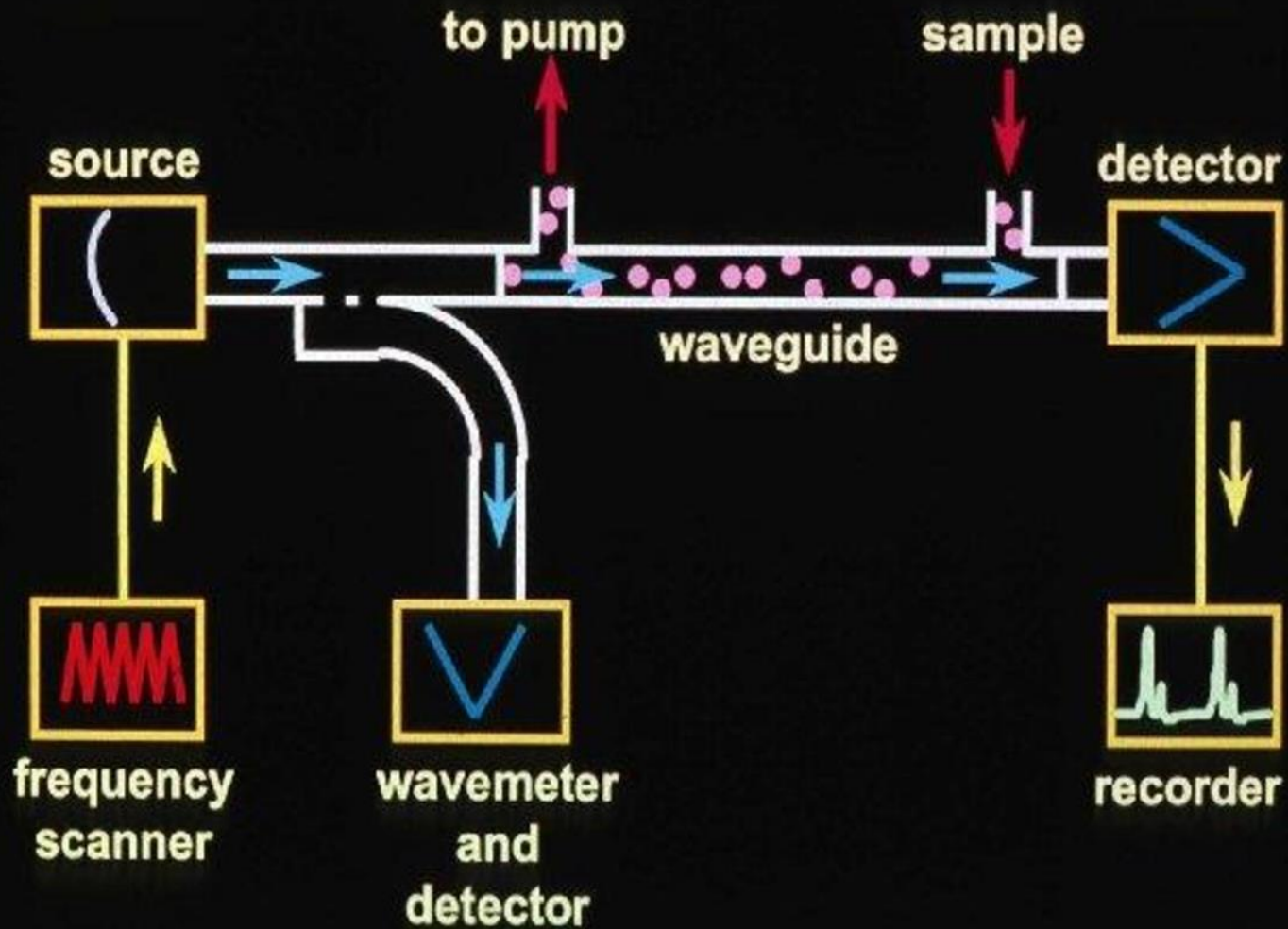
Microwave Spectroscopy

- RADAR was impetus for its invention
- 1948, Walter Gordy, first published review

Microwaves



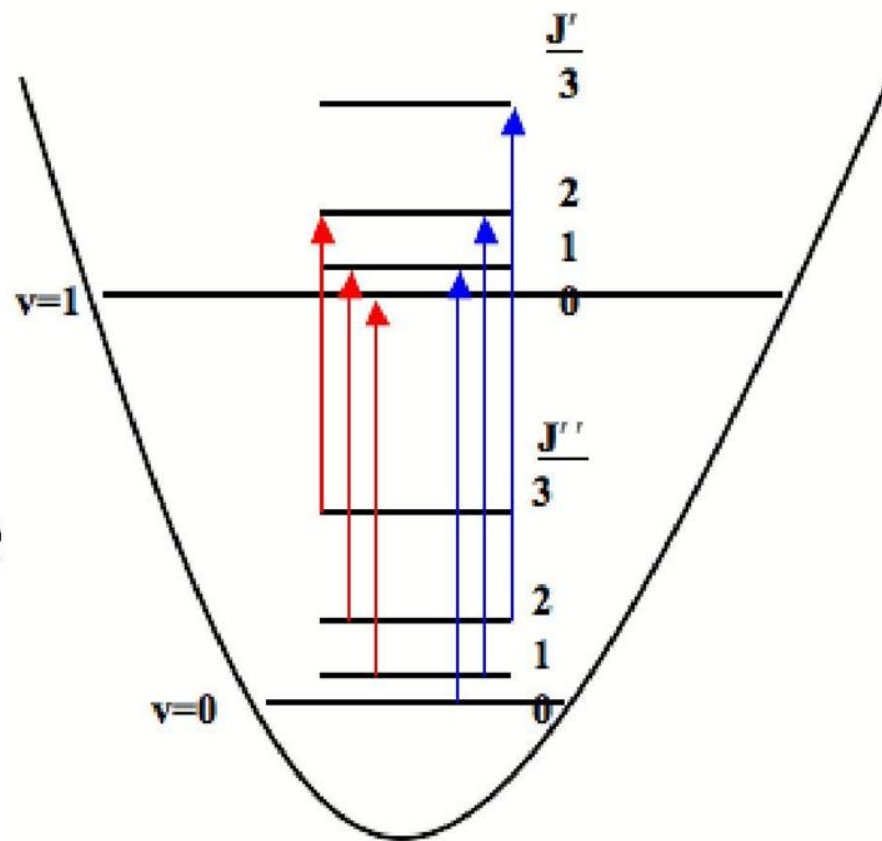
Microwave spectrometer



Rotational vs. Vibrational

- Where J = rotational
– V = vibrational
- Observed from lowest vibration state

$$E(J) = B J(J + 1)$$

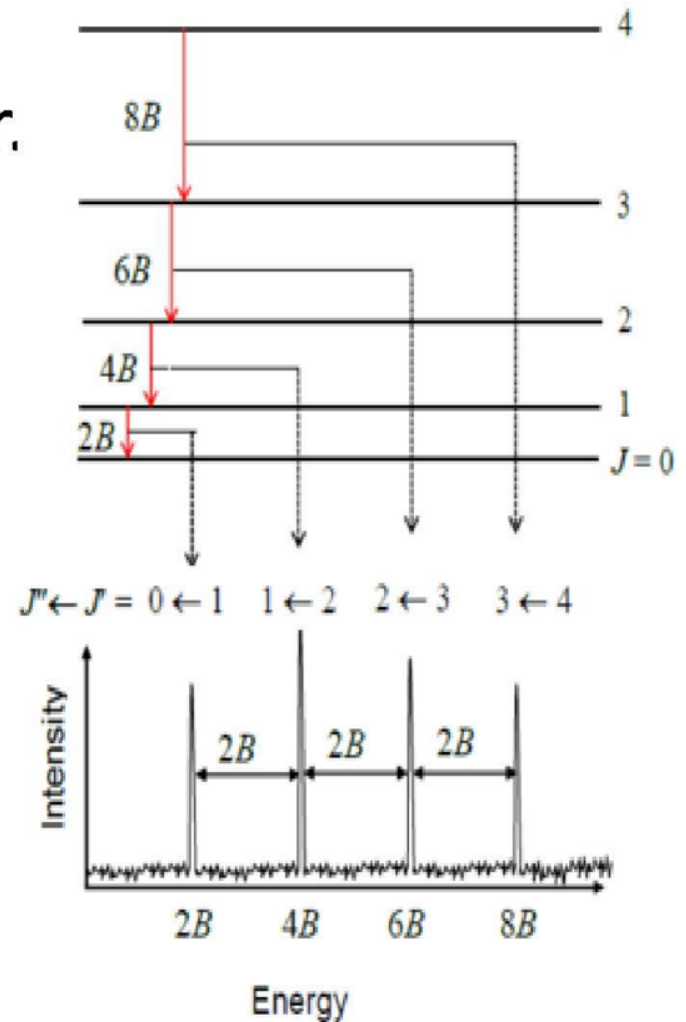


Motion of Rotation

- 3 Possible moments of inertia
 - I_a, I_b, I_c

Diatomic molecule

- Rotation E from Schrodinger.
- $I_a = I_b, I_c = 0$
- $B = \frac{h}{8Ic\pi^2}$
- $I = R^2 u = \frac{(M_1 M_2)}{(M_1 + M_2)} R^2$
- Line spacing is $2B$.



Theory of Microwave Spectroscopy

- Microwave wavelength photon
- Highest probability of transition
- Molecules with dipoles

How do we measure if it will translate?

- Probability of Transition = $\int \psi_{\text{rot}}^*(F) \mu^{\wedge} \psi_{\text{rot}}(I) d\tau$
 - Where: $\psi_{\text{rot}}^*(F)$ is the complex conjugate of the final rotational state
 - $\psi_{\text{rot}}(I)$ is the wave function of the initial rotational state
 - μ is the dipole moment operator with X, y, z coordinates.
- The function is positive.
- Only tells if is allowed.

But there are limits.

- Photons limited
 - Each photon has one unit of momentum

$$\Delta J \pm 1$$

– Only one transition per

Is it available to translate?

- Boltzmann distribution

$$\frac{n_J}{n_0} = e\left(-\frac{E_{rot}(J)}{RT}\right) / \sum_{J=0}^{J=n} e\left(-\frac{E_{rot}(J)}{RT}\right)$$

Where n_J =number of molecules excited

n_0 =number of molecules in ground state

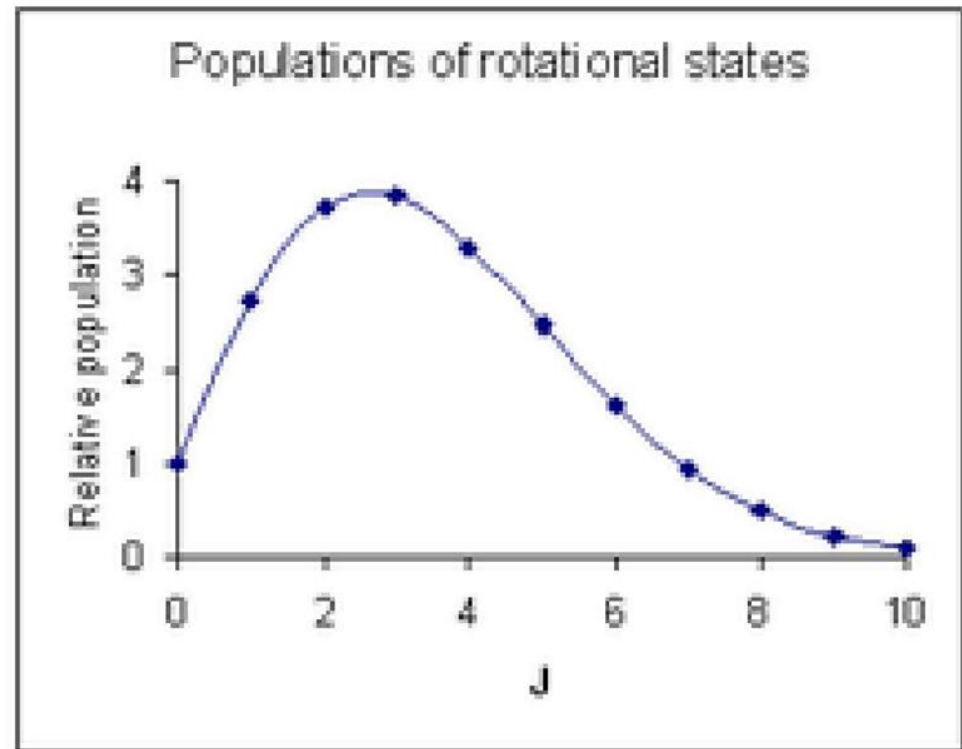
R =gas constant

T =temperature

$E_{rot}J$ =molar energy of the rotational state

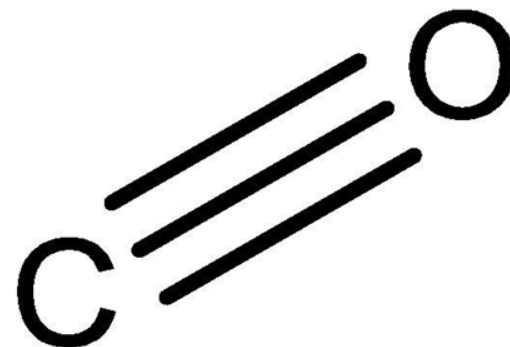
Plot of Probability

- Probability of Population
- Similar to spectra



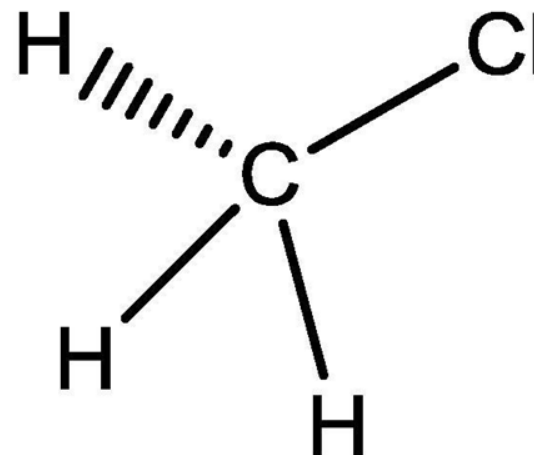
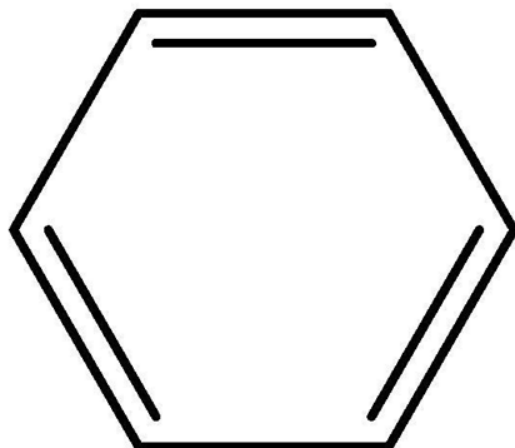
Types of Molecules that are MW-Active

- Linear



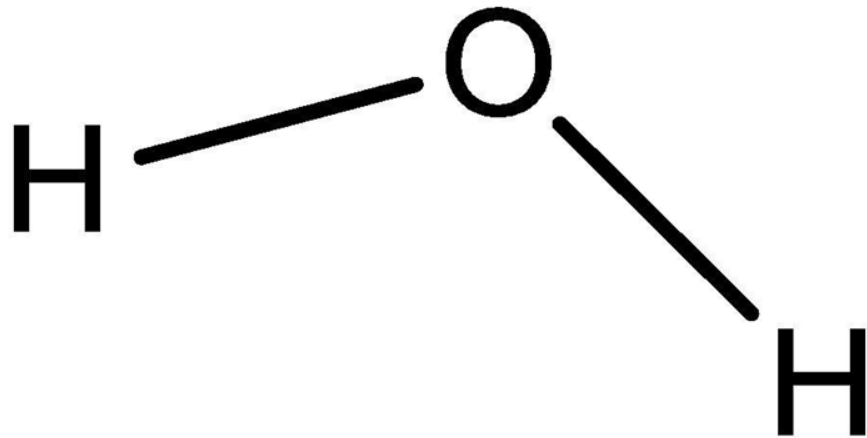
Types of Molecules that are MW-Active

- Symmetric Tops



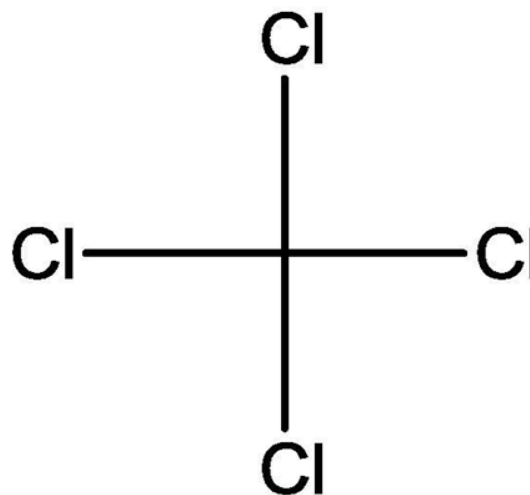
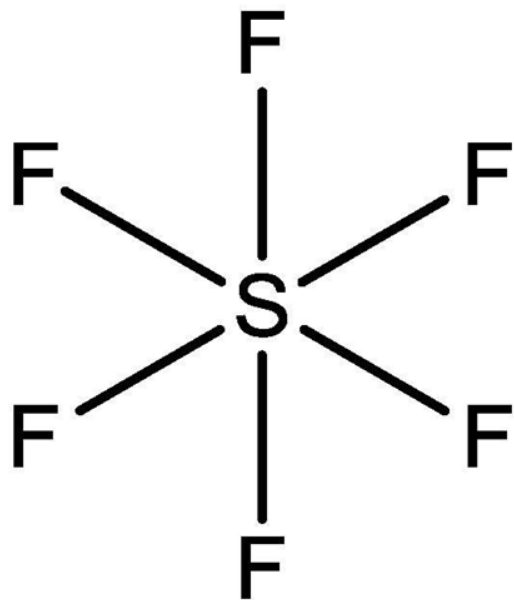
Types of Molecules that are MW-Active

- Antisymmetric tops



Types of Molecules that are MW-Active

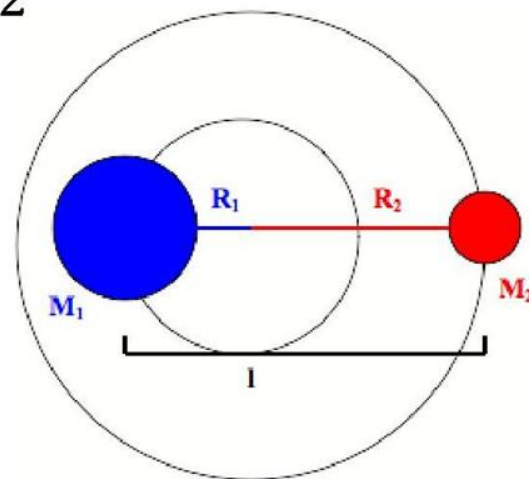
- Spherical tops-*Not active*



Linear Molecules.

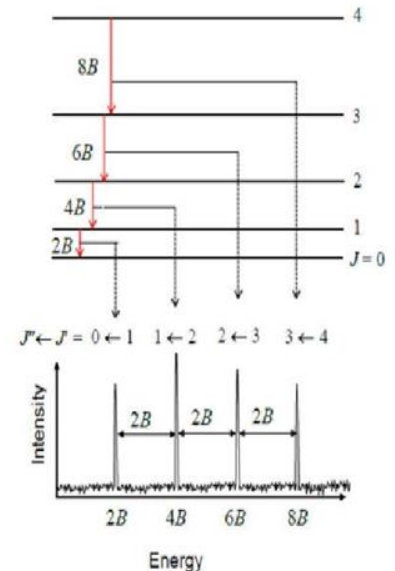
- Linear molecules(Rigid rotors)
- Bond length directly calculated
- $I_a = I_b , I_c = 0$

$$I = R^2 u = \frac{(M_1 M_2)}{(M_1 + M_2)} R^2$$



Linear Molecules.

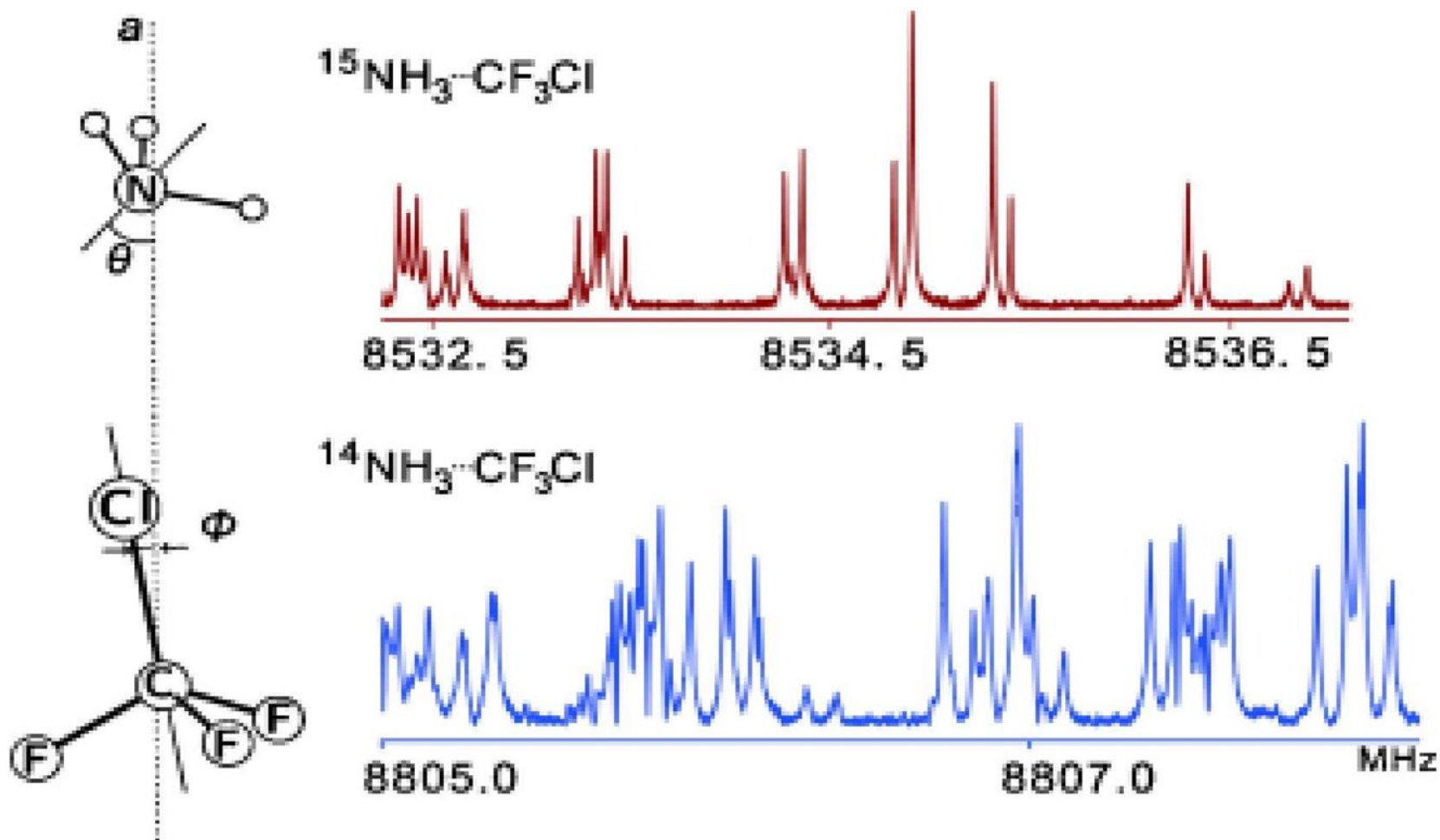
- Linear molecules(Rigid rotors)
- $\Delta J = J' - J'' = \pm 1$
- $\nu_{J' \leftrightarrow J''} = E(J') - E(J'') = 2B(J'' + 1)$
 - Where $J'' = \text{Lower level}$
 - $J' = \text{Upper Level}$



Symmetric Tops (Rotors)

- Two rotational axes same, one different
- Oblate
 - $I_a \neq I_b = I_c$
 - Oblate.
 - Benzene, XeF₄
- Prolate
 - $I_a = I_b \neq I_c$
 - Prolate.
 - CH₃Cl, NH₃

Symmetric Top Spectra



Fengm G., Evangelisti, L, Gaspartini, N., Caminati, W., *On the Cl...N Halogen Bond: A Rotational Study of $\text{CF}_3\text{Cl}\cdots\text{NH}_3$* , Chemistry-A European Journal, Vol 18, #5 pg 1364-1368

Why more complicated?

- With more axis, more complicated
 - Symmetrical molecules gain 2 terms
 - K = vector about the symmetry axis
 - Must be between $-J$ and $+J$
 - M = rotational momentum about a external field
 - Also between $-J$ and $+J$
 - 0 if no external field

Why more complicated?

- With more axis, more complicated

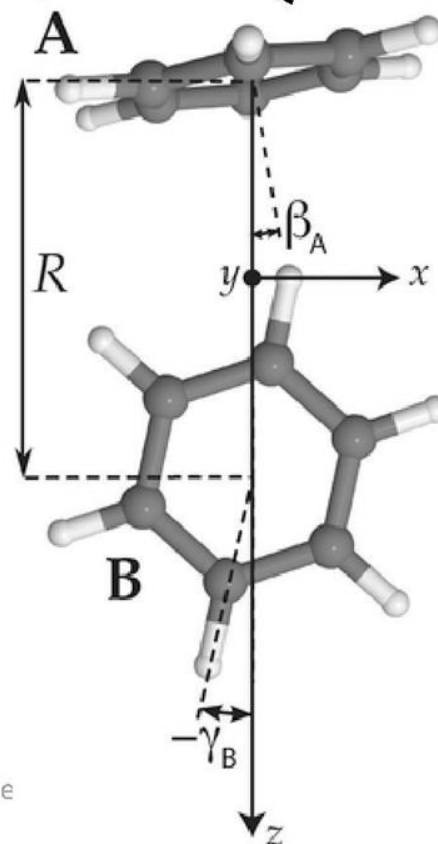
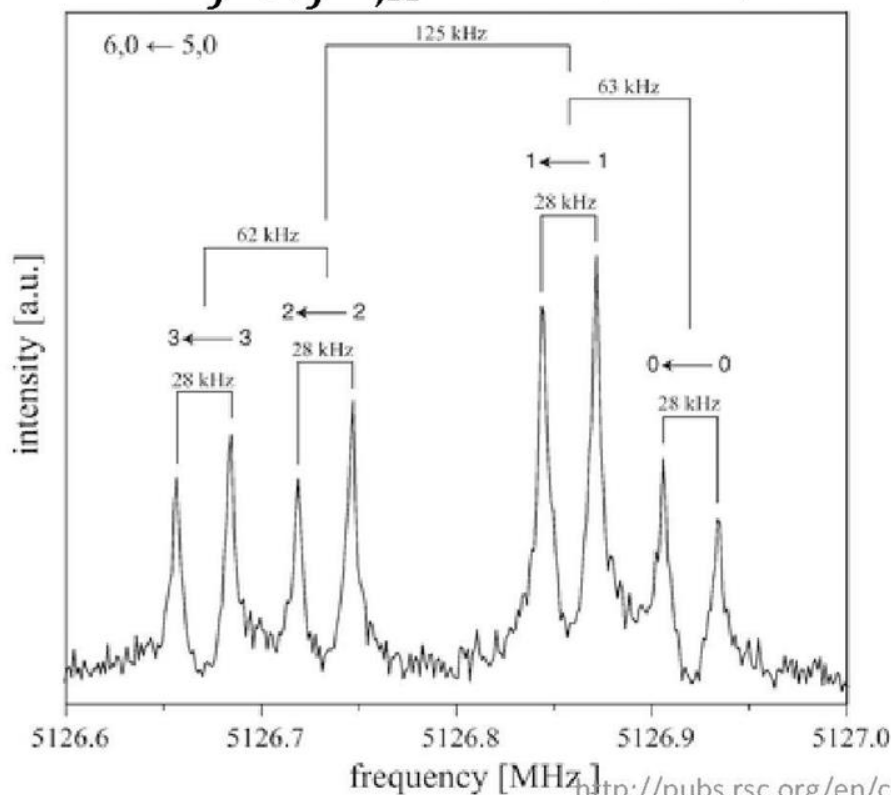
$$- E(J, K) = BJ(J + 1) + (A - B)K^2$$

- Where $B = \frac{h}{8\pi^2 c I_b}$
- $A = \frac{h}{8\pi^2 c I_a}$, Prolate
- $A = \frac{h}{8\pi^2 c I_c}$, Oblate

Why more complicated?

- Which leads to Lines at:

$$-\nu_{J' \leftrightarrow J'', K} = E(J', K) - E(J'', K) = 2B(J'' + 1)$$



Stark Effect

- Similar to Zeeman effect
- Lifts level degeneracy
- Due to external electric field
- 1st Order-Linear
- 2nd Order-Quadratic

Stark Effect

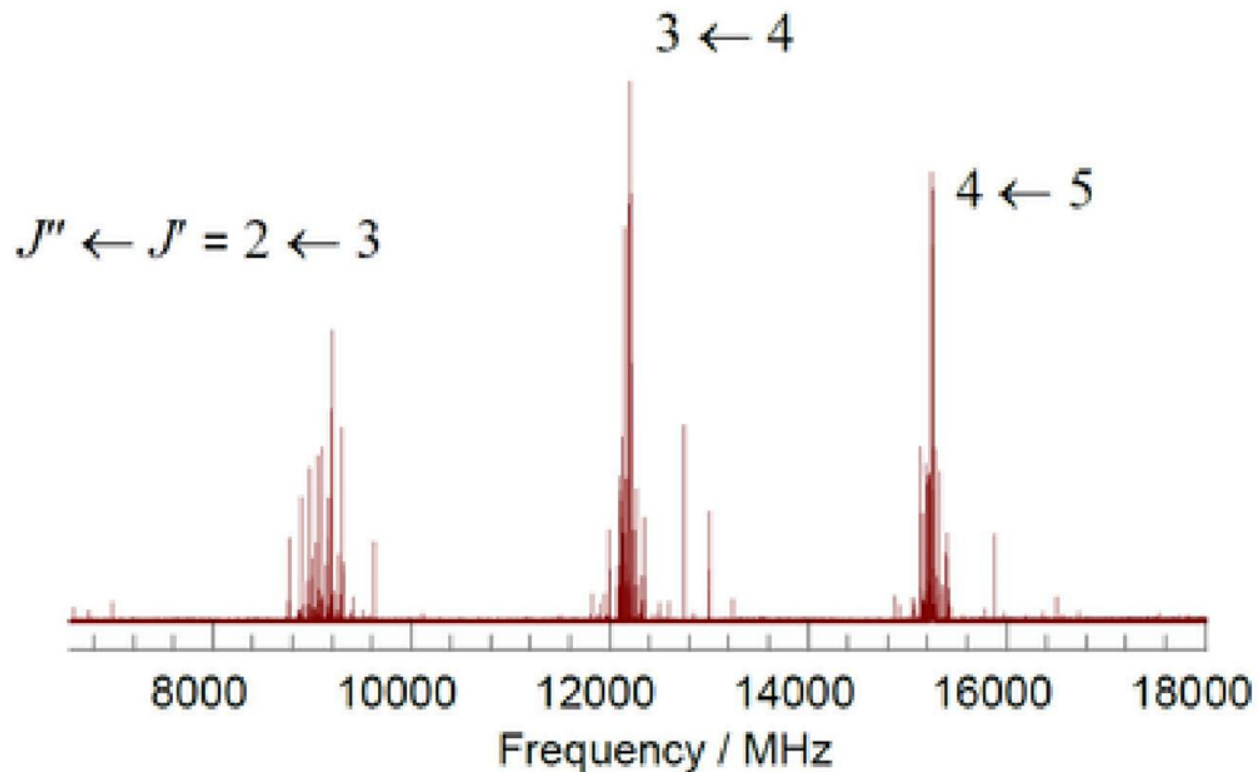
- $E_{Stark}(J, K, M) = -\frac{uEKM}{J(J+1)}$
- $M = 0$, no stark
- $K, M \neq 0$ splitting occurs
 - $-M$ inc, $+M$ dec

Hyperfine Splitting

- Coupling of Nuclear spin and molecular rotation.
 - If $J > I$, $2I+1$ levels
 - If $J < I$, $2J+1$ levels.

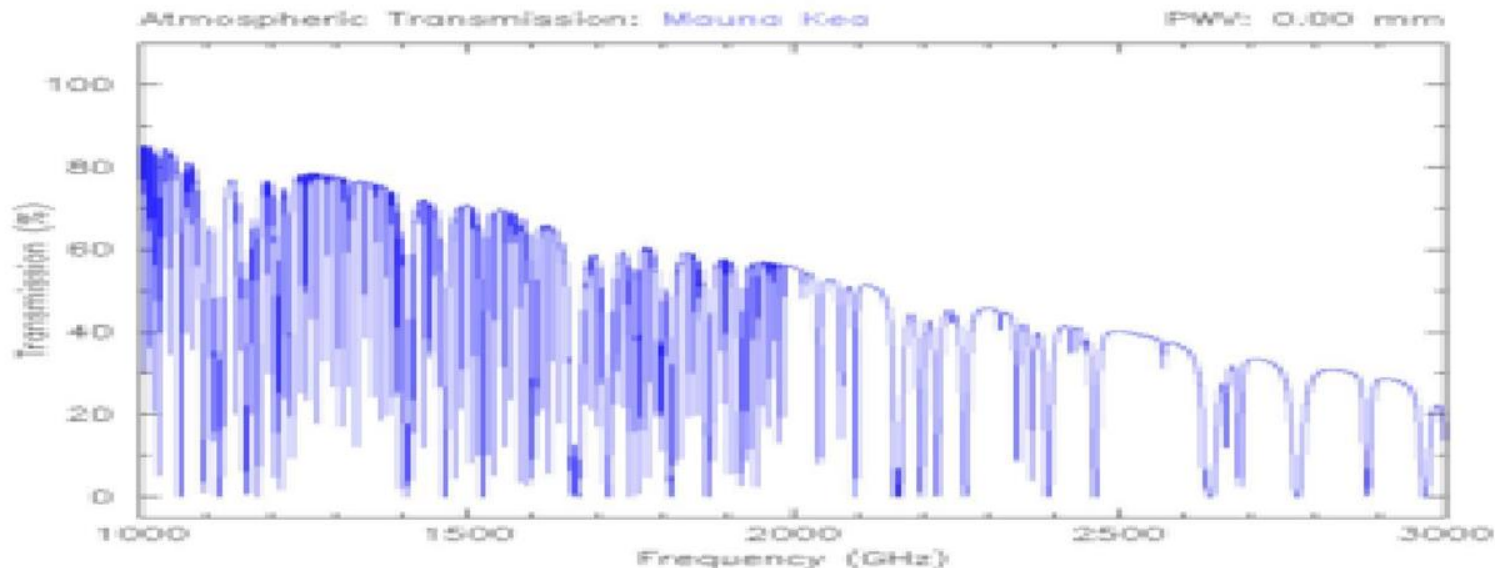
Hyperfine splitting

- CF_3I , Splitting due to ^{127}I



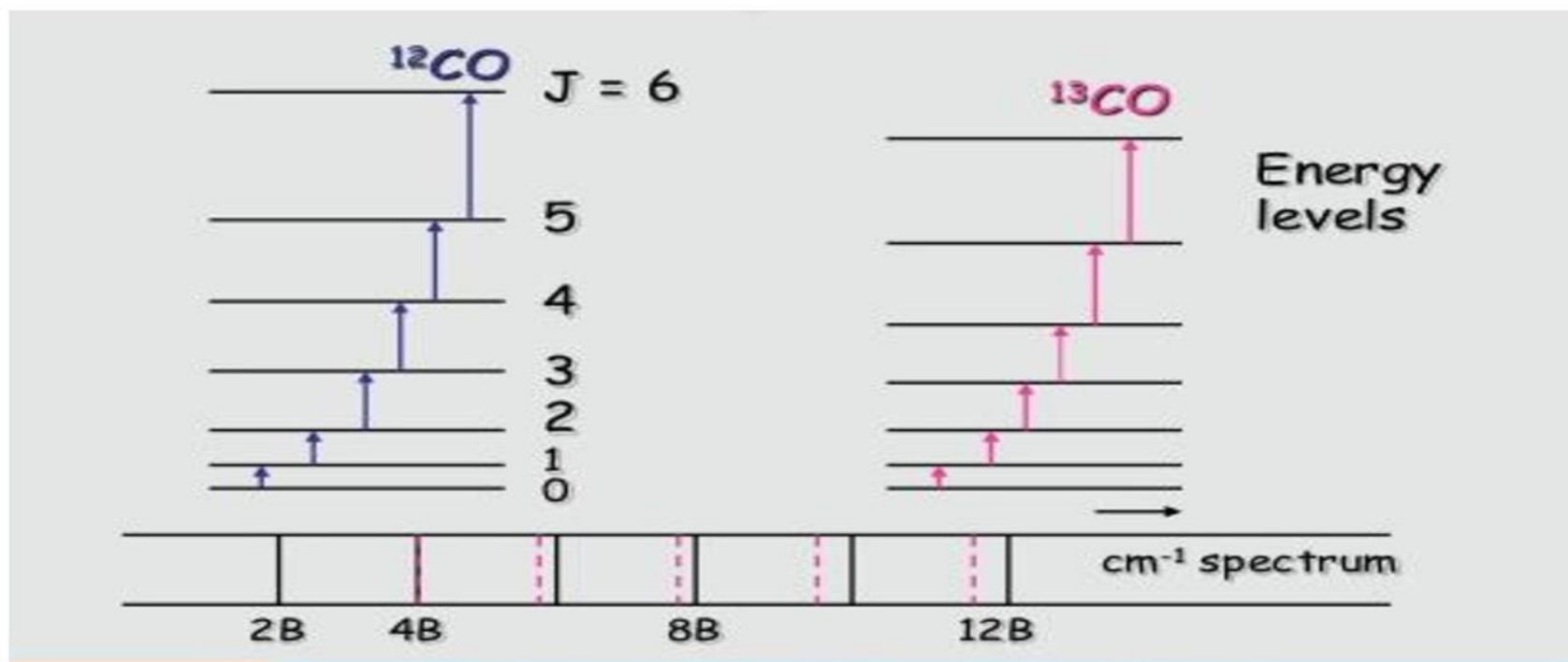
Asymmetrical Tops

- Three different axes, 3 different inertias
- Most molecules.
- Very complex spectra.



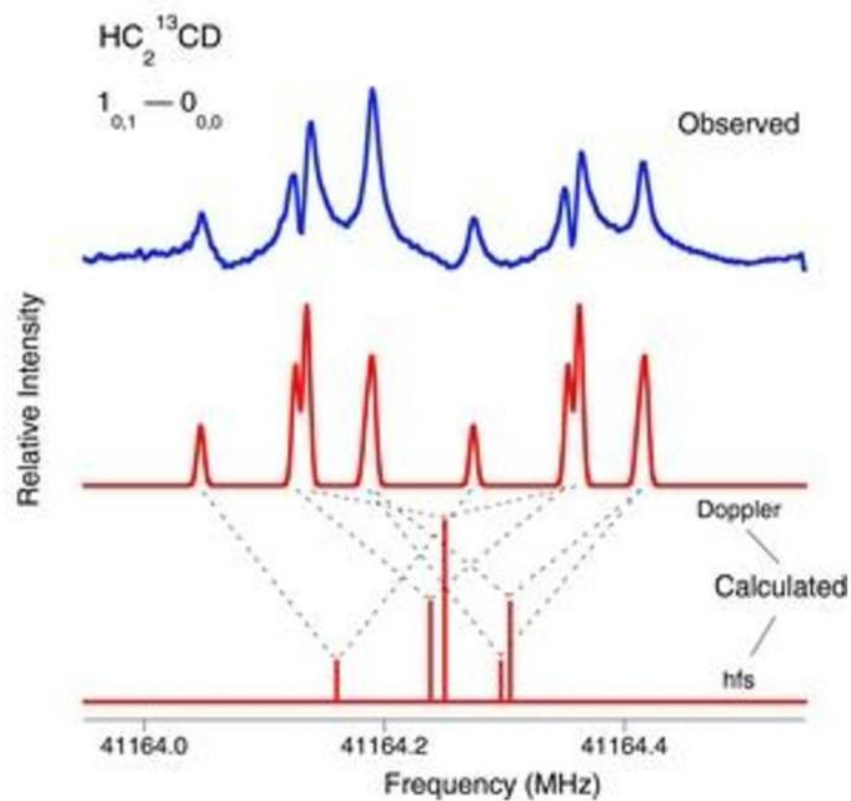
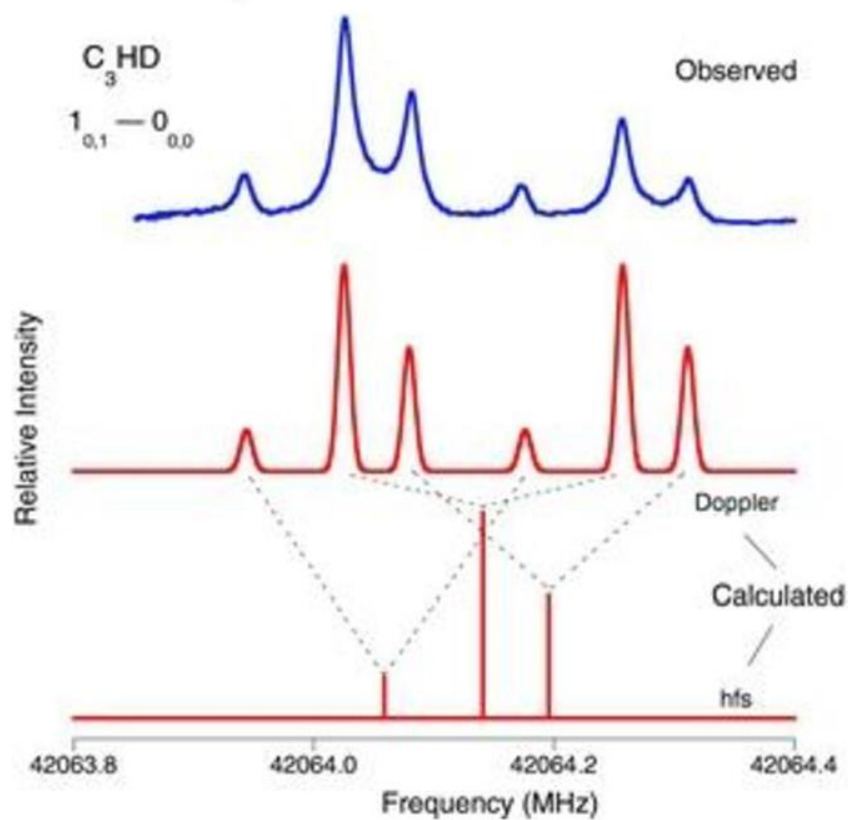
Effect of Isotope Substitution

- $^{12}\text{C}^{16}\text{O} \rightarrow ^{13}\text{C}^{16}\text{O}$, mass \uparrow , $B \downarrow$ ($\sim 1/I$), $E \downarrow$.



Isotope splitting

- $c\text{-C}_3\text{HD}$**



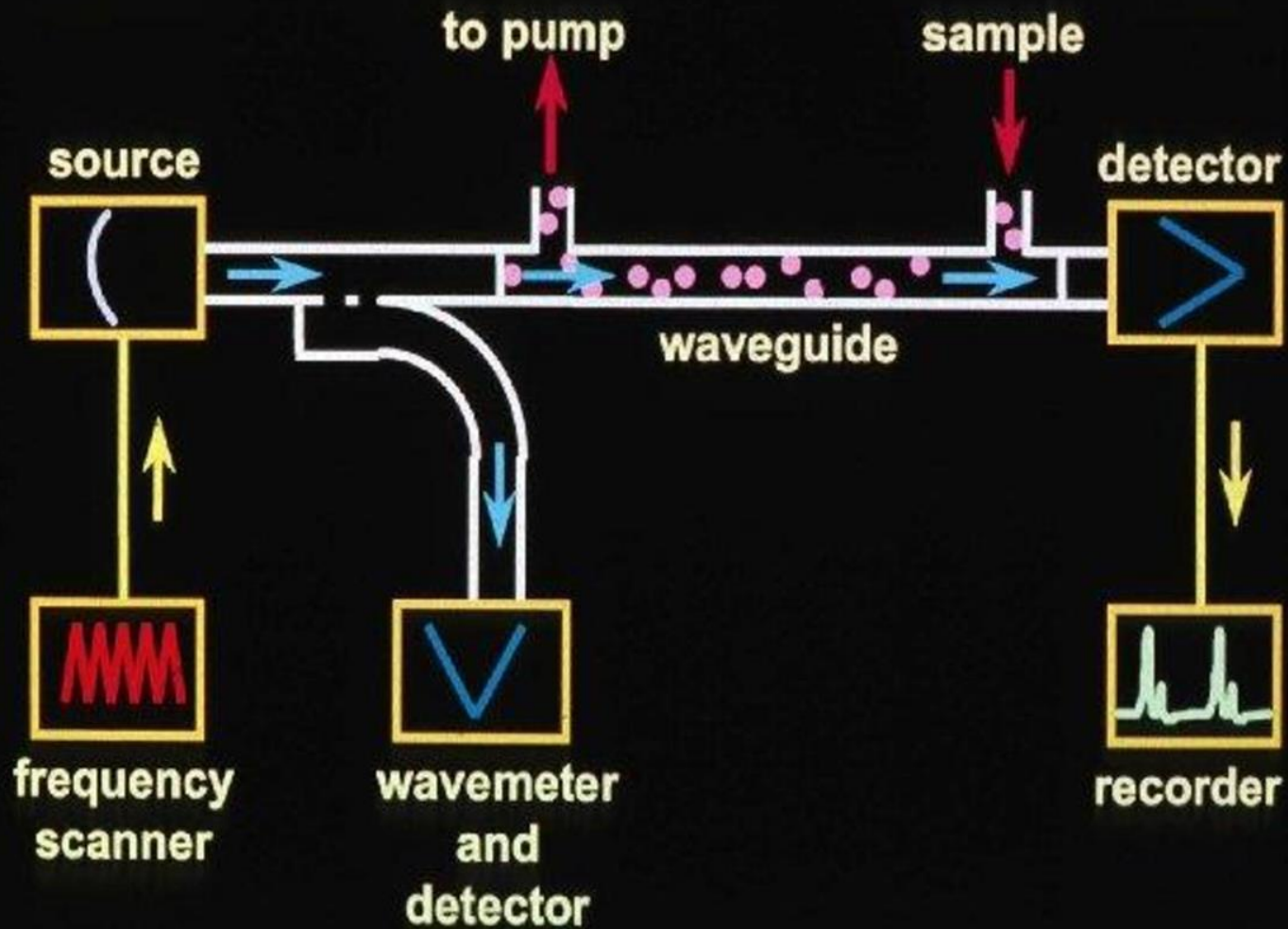
Instrumentation

- Most homemade.
- Two types spectrometers.
- Stark Modulated
- Fortier Transform Microwave Spectroscopy
 - Similar in concept to FTIR

Stark Modulated MW Spectrometer

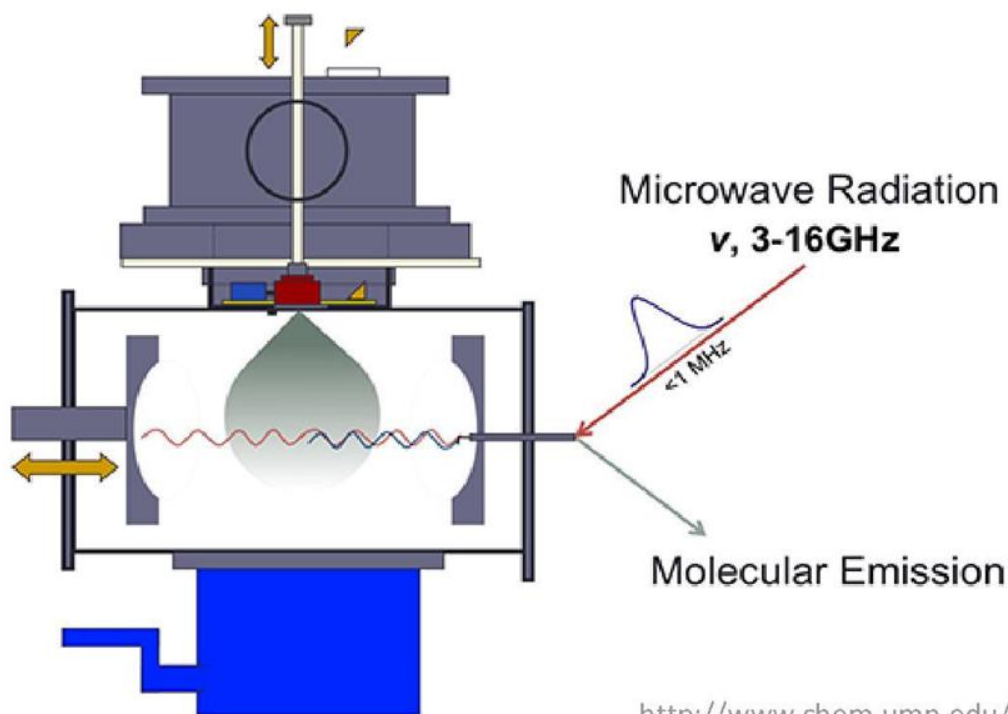
- Samples introduced as a gas.
- Can be heated.
- Generally high vacuum.

Microwave spectrometer



FTMW Spectrometer

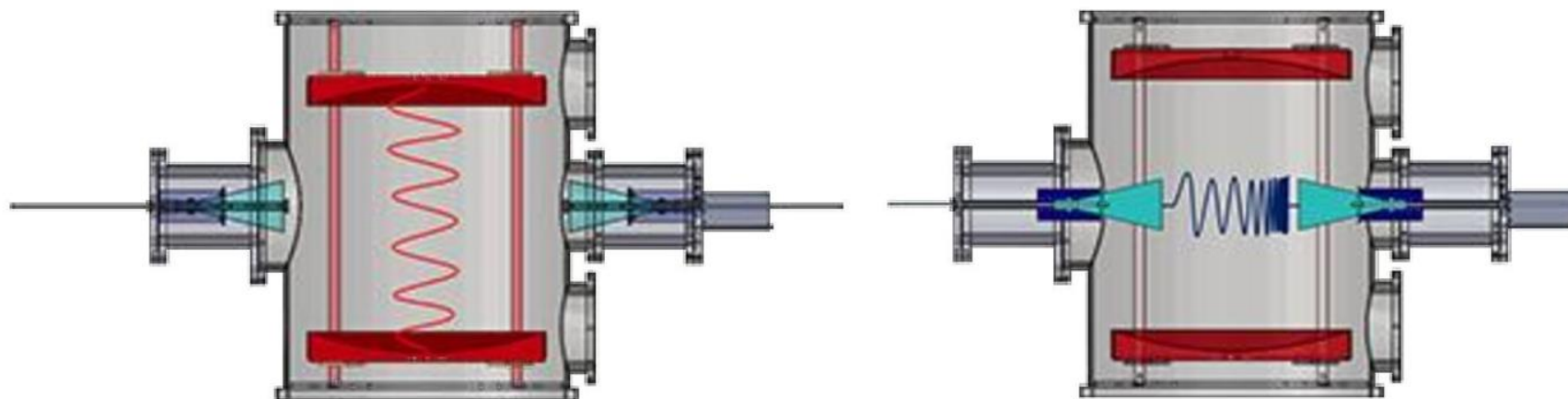
- Similar in principle to FTIR
- Broader Frequency, Greater Precision



Chipped Pulse FTMW

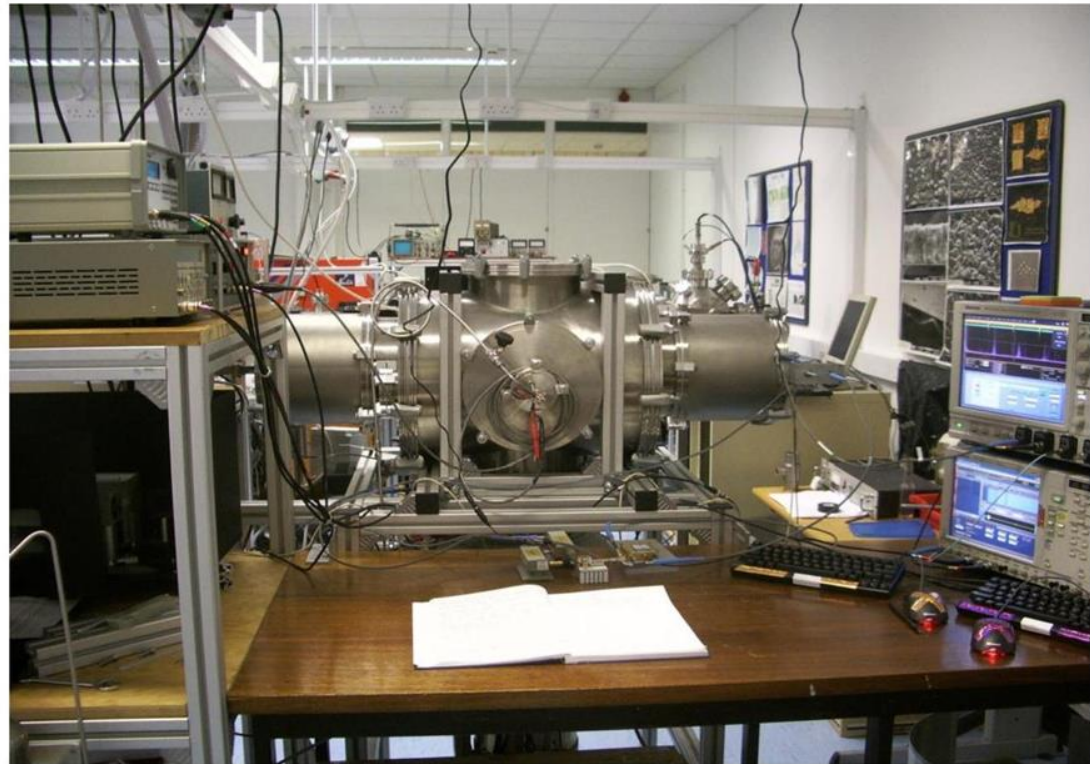
- Add waveform generators
- Widens Bandwidth several 1000x
- Decreases spectral search time

Chipped Pulse FTMW



FTMW Spectrometer

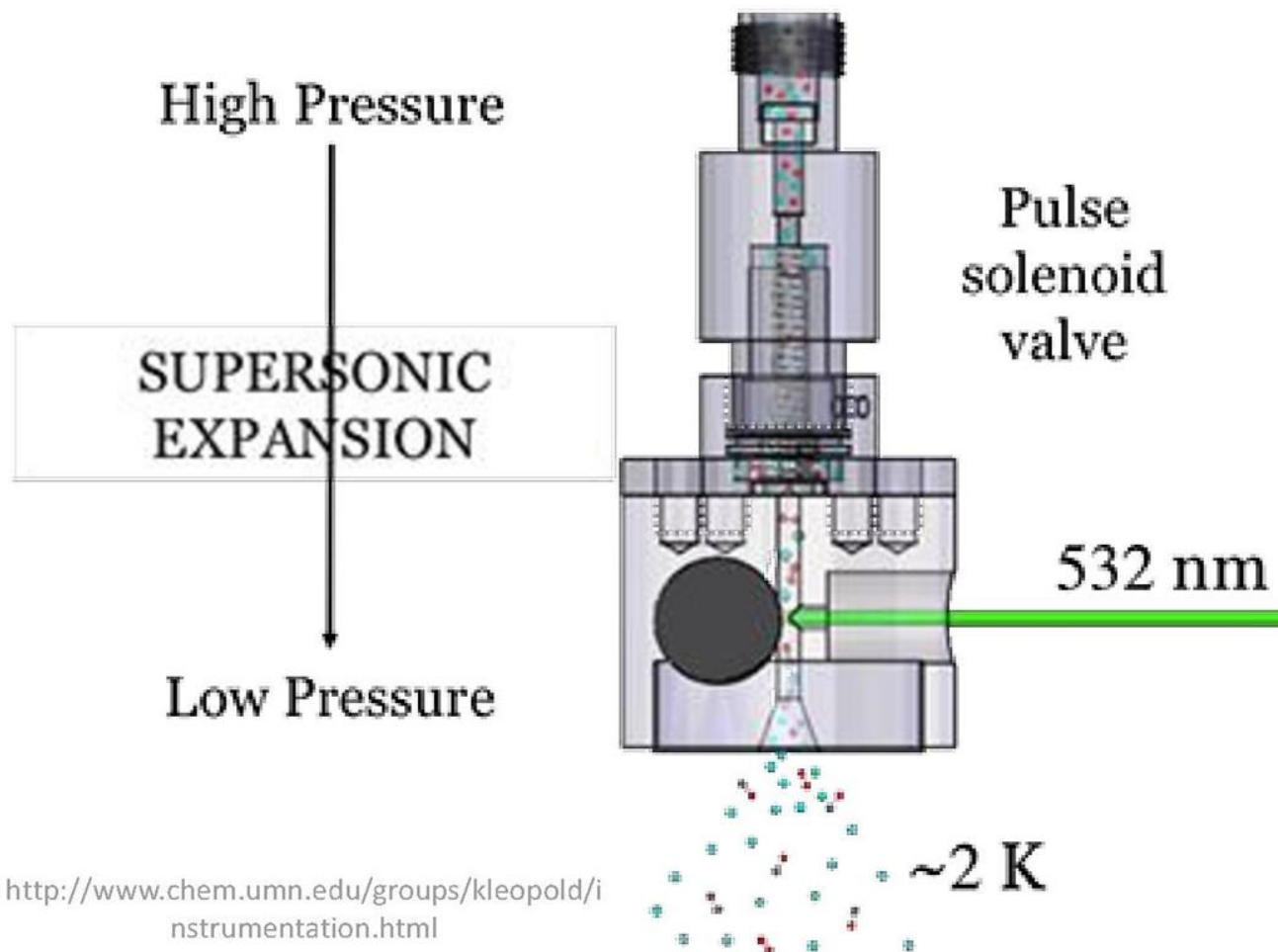
- Typical setup



Sample Preparation/Introduction

- Gas samples
 - As is
 - As a diluted analyte in a non-MW reactive gas such as Neon
- Solids
 - Laser abatement
 - Vaporization
- Liquids
 - Vaporization
 - Supersonic Expansion

Supersonic Expansion/Laser Ablatement



Applications of MW

- Measurement of bond lengths
- Observation by radio telescopes for life precursors in interstellar clouds
- Precise observation of translating stereochemistries and confirmation verification

Measuring Bond Length.

- For example, we will use the easiest case, a diatomic molecule, HCL.

| cm ⁻¹ | J->J+1 | R(nm) |
|------------------|--------|-------|
| 83.03 | 3-4 | .1288 |
| 103.73 | 4-5 | .1288 |
| 124.3 | 5-6 | .1289 |
| 145.03 | 6-7 | .1289 |
| 165.51 | 7-8 | .1290 |
| 185.86 | 8-9 | .1291 |
| 206.38 | 9-10 | .1292 |
| 226.5 | 10-11 | .1293 |

Calculation example (HCL)

$$B = \frac{h}{8Ic\pi^2}$$

$$R = \sqrt{\frac{h}{8\pi^2 c B u}}$$

R=

$$\sqrt{\frac{6.626e-34 \text{ J.S}}{8(3.14)^2 (2.99792458e10 \text{ cm.s}) \left(\frac{35.5}{36.5} * 1.661 \times 10^{-27} \text{ kg}\right) (10.3 \text{ cm}^{-1})}}$$

$$R = 1.29 \times 10^{-10} \text{ M} \text{ or } 0.129 \text{ nm}$$

Calculation example (CO)

$$B = \frac{h}{8Ic\pi^2}$$

$$R = \sqrt{\frac{h}{8\pi^2 c B u}}$$

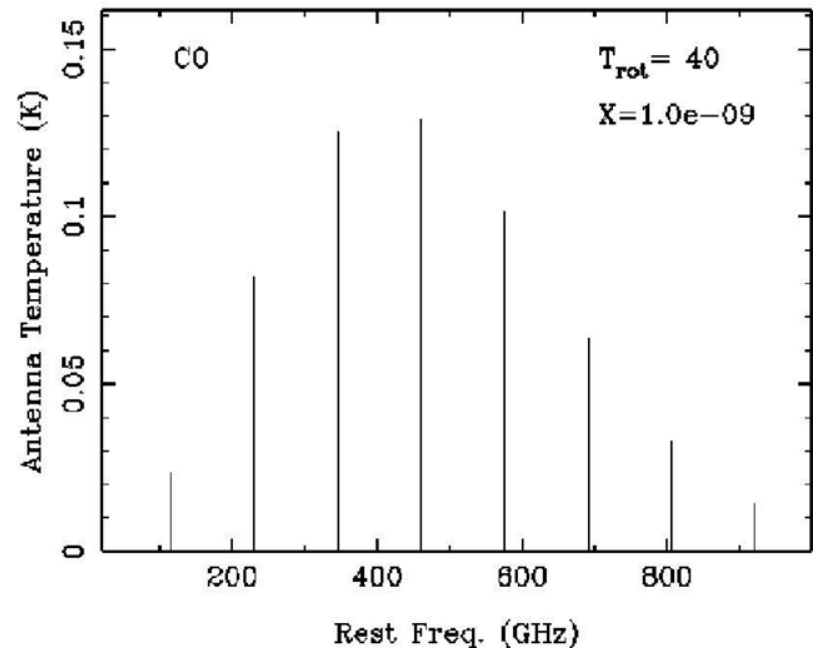
R=

$$\sqrt{\frac{6.626e-34 \text{ J.S}}{8(3.14)^2 (2.99792458e10 \text{ cm.s})\left(\frac{48}{7}\right)*1.661x10-27 \text{ kg})(1.9313 \text{ cm}^{-1})}}$$

$$R=1.13x10^{-10} \text{ M or } 0.113 \text{ nm}$$

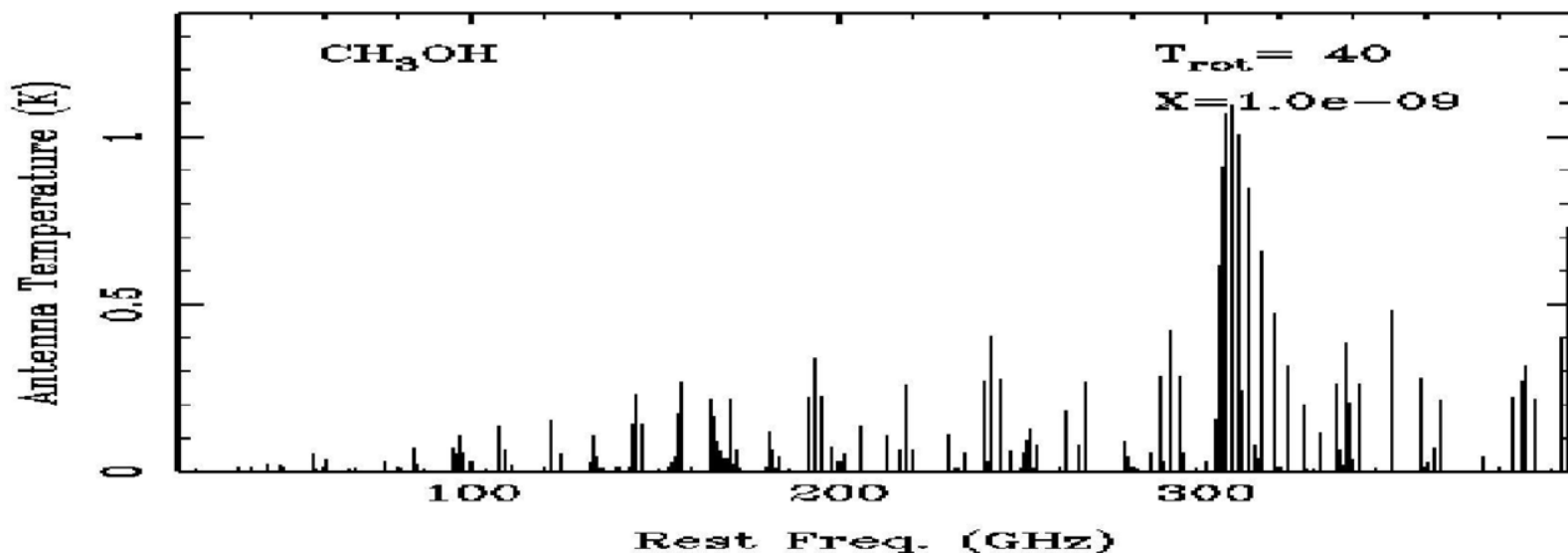
Identification of Organics in Interstellar Space

- Gas phase molecules and radicals
- Observed using a radio telescope.
- Linear molecule
 - Simple spectra



Identification of Organics in Interstellar Space

- Symmetric Molecule
 - Complicated due to symmetric top
 - Many more possible states



Conclusion

- Useful for Gas Molecules
- Can determine bond length.
- Diverse usages for the technique.